# EXPRESSIONS OF THE ANISOTROPY OF THE CRUST BY MEANS OF THE OBSERVATIONS OF THE EARTH TIDAL STRAINS 

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#### Abstract

A weighted function between an azimuthal pattern of an observed tidal strain and that of a theoretically calculated tidal strain is obtained in order to express an anisotropy of the crust.

The weighted functions are calculated for the azimuthal pattern of the tidal strains at Osakayama, Kishu and Suhara. The phase angles of the weighted functions show that the axes of the anisotropy at Osakayama concord with the normal direction of the Japan Proper, and the axes of the anisotropy at Kishu concord with the axis of Kii Peninsular. The load tide is calculated as the difference between the observed tidal strain at Suhara and the theoretical direct effect of the earth tide for the earth's model of Gutenberg. The weighted function between this load tide obtained by the observation and the theoretical load tide calculated by use of the oceanic tide is calculated. This weighted function shows that the axis of the load tide is deffected by about $45^{\circ}$ by the metamorphical geologic constructions around Suhara district.


## 1. Introduction

Everyone believes the earth crust has an inhomogeneous constitution. For example, information on the distributions of the seismic velocity in the crust shows that the earth crust is an amalgam of materials which have various elastic constants. It is very hard to estimate by means of how to amalgamate the different materials in the crust by using only the seismic wave velocity. It is the reason why the estimation of the behavior of the whole earth is difficult that even an anomalous materials which occupy very thin layers of the earth are importance. So, it is important to study the behavior of the whole earth as a phenomenon of the earth's tidal strain.

The earth tidal strain consists of the direct effect due to the tide generating force of the moon or the sun, and the in-direct effect due to the oceanic tide. The in-direct effect at near ocean is much important as same as the direct effect. H. Takeuchi et.al. [Takeuchi (1965)] and I. M. Longman [Longman (1963)] have calculated the direct and in-direct effect of the tidal strains of the model earth of B. Gutenberg. It, however, is difficult to calculate the in-direct effect (load tide) at any place. As 5/7 of the earth's surface is covered with ocean, it is unnatural that the in-direct effect is isolated from
the effect on the whole earth.
The author tries to obtain the in-direct effect at any point as a deflection from the mean value of the whole earth or the theoretically estimated values of the effect for the model earth. He also treats the deflection as the isotropy of the crust. Now, this is meant the crust which loads the ocean.

## 2. Principle

a) Weighted function. C. Tsuboi [Tsuboi (1940)] has explained concisely the weighted function or the means, and applications for the relations between the gravity anomaly and the mechanism of the isostasy. The calculations and the interpretations of the weighted function are complex in a general case, because the function is applied for nonperiodic functions. Naturally, the weighted function should be applied for the calculation of the relation between periodic functions.

An expression of the azimuthal pattern of the linear strain is given a function

$$
\begin{align*}
F(p) & =A+B \cos 2 p+C \sin 2 p \\
& =A+\sqrt{B^{2}+C^{2}} \cos \left(2 p-\tan ^{-1} \frac{C}{B}\right) . \tag{1}
\end{align*}
$$

This function $F(p)$ has only 3 constants $A, B$ and $C$ or $A, \sqrt{B^{2}+C^{2}}$ and $\tan ^{-1}$ $\frac{C}{B}$.

The applied force of $F(p)$ which is the observed earth-tidal strain is the tidegenerating force. Let $f(p)$ be the azimuthal function of the theoretical function of the direct of the earth tide for the model earth as follow,

$$
\begin{equation*}
f(p)=a+b \cos 2 p+c \sin 2 p \tag{2}
\end{equation*}
$$

$F(p)$ and $f(p)$ are combined with a weight function shown as follows,

$$
F(p)=\int_{0}^{2 \pi} f(p+\tau) \varphi(\tau) d \tau
$$

where

$$
\begin{align*}
& \varphi(\tau)=\alpha_{0}+\sqrt{\alpha_{2}^{2}+\beta_{2}^{2}} \cos \left(2 \tau-\tan ^{-1} \frac{\beta_{2}}{\alpha_{2}}\right), \\
& \alpha_{0}=\frac{1}{2 \pi} \frac{A}{a}, \quad \alpha_{2}=\frac{1}{\pi} \frac{a A+b B}{a^{2}+b^{2}}, \quad \beta_{2}=\frac{1}{\pi} \frac{b A-a B}{a^{2}+b^{2}} \tag{3}
\end{align*}
$$

b) Direct effects of the earth tide. The deformation of the spheroidal type only is usually considered in the earth tide. The horizontal strain elements, $e_{\theta 0}, e_{\phi \phi}$ and $e_{\theta \phi}$, of the spheroidal type deformation of the semi-diurnal and diurnal components are given as follows

$$
\left.\begin{array}{rl}
e_{\theta \theta} & =\sum_{i}\left(h_{2}-4 l_{2}\right) \frac{W_{2}(1)_{i}}{a g}+\sum_{i} \frac{h_{2} \sin ^{2} \theta+2 l_{2} \cos 2 \theta}{\sin ^{2} \theta} \frac{W_{2}\left(\frac{1}{2}\right)_{i}}{a g} \\
e_{\phi \phi} & =\sum_{i}\left(h_{2}-2 l_{2}\right) \frac{W_{2}(1)_{i}}{a g}+\sum_{i} \frac{h_{2} \sin ^{2} \theta-2 l_{2}\left(1+\sin ^{2} \theta\right)}{\sin ^{2} \theta} \frac{W_{2}\left(\frac{1}{2}\right)_{i}}{a g} \\
e_{\theta \phi} & =\sum_{i} 4 l_{2} \sin \theta \cdot \tan (t+\phi) \frac{W_{2}(1)_{i}}{a g}  \tag{4}\\
& -\sum_{i} 4 l_{2} \cos \theta \cdot \tan 2(t+\phi) \frac{W_{2}\left(\frac{1}{2}\right)_{i}}{a g}
\end{array}\right\}
$$

where $W_{2}(1)_{i}=J_{1 i} \sin 2 \theta \cos (t+\phi)$, and is the diurnal component of the potential of the tide generating force. $W_{2}\left(\frac{1}{2}\right)_{i}=J_{2 i} \sin ^{2} \theta \cos 2(t+\phi)$, and is the semi-diurnal component of the potential of the tide generating force. $J_{1^{i}}$ and $J_{2^{i}}$ are the special constants of the component tides of the diurnal and semi-diurnal components, respectively. $h_{2}, l_{2}, a, g, \theta, \phi$ and t are Love's number, Shida's number, the mean value of the acceleration of the gravity at the observatory, the colatitude, the east longitude and the hour angle of the heavenly body, respectively. According to formulas (4), these strain elements are decided by the values of $h_{2}$ and $l_{2}$. For example, the phase angles of $e_{\theta \theta}$ and $e_{\phi \phi}$ are $0^{\circ}$ in the case of that $h_{2}=0.612$, and $l_{2}=0.083$ which are calculated by H. Takeuchi for the model earth of B. Gutenberg. I. M. Longman has obtained also as nearly equal values. However, the phase angles of $e_{\theta \theta}$ in diurnal and of $e_{\phi \phi}$ in semi-diurnal are $180^{\circ}$ in the case $h_{2}=0.600$ and $l_{2}=0.200$ which shows the incompressible earth model. Now, we should examine the values of $h_{2}$ and $l_{2}$ again. According to the theoretical calculation, the horizontal areal strain $e_{\theta \theta}+e_{\phi \phi}$ is free from the effect of the load tide, and the phase angle of the observed areal strain is also almost equal to $0^{\circ}$. And the amplitude of the horizontal areal strain obtained by observation distant from the occean is always greater than both of $e_{\theta \theta}$ and $e_{\phi \phi}$. Therefor, we can estimate $l_{2}$ is much smaller than $h_{2}$, and $I_{2}$ is smaller than $h_{2} / 3$ at least. $l_{2}$ can be said to be much smaller than $0.204 h_{2}$, because the observed phase of $e_{\theta \theta}$ of $0_{1}$-tide is nearly equal to $0^{\circ}$, and the observed amplitude is large considerably. Because the observed phase of the radial component $e_{r r}$ of the strain is nearly equal to $180^{\circ}$, we can also estimate that the value of $a \frac{d h_{2}}{d r}$ near the earth's surface are considerably larger than $2 h_{2}$. That shows $h_{2}$ decreased remarkably near the earth's surface.

## 3. Observations

We have three observatroies [I. Ozawa, (1957, 1966)], Osakayama ( $34^{\circ} 59.6^{\prime}$ of the north latitude, $135^{\circ} 51.5^{\prime}$ of the east longitude), Kishu ( $33^{\circ} 51.7^{\prime}$ of the north latitude, $135^{\circ} 53.4^{\prime}$ of the east longitude) and Suhara ( $34^{\circ} 02.6^{\prime}$ of the north latitude,
$135^{\circ} 11.7^{\prime}$ of the east longitude) in Kinki District of Japan. The positions of these observatories are shown in Fig. 1. Their distance from sea is 65 km from Osakayama, is 16.5 km from Kishu, and is about 100 m from observing room at Suhara. All observations are conducted under the ground surface at depths more than 50 m or 100 m . The observed tidal constants at these observatories are listed in Table 1.

The observed tidal strain elements and their horizontal main strains are calculated from the Table 1, and are shown in Table 2, 3, 4 and 5. The strain elements are calculated also in Tables 2,3 and 4 in both the cases that $h_{2}=0.612$ and $l_{2}=0.083$, and that $h_{2}=0.600$ and $l_{2}=0.200$. And the elements and the main strains of some load tide of $M_{2}$-components of the oceanic tides at these three observatory and $0_{1}$-component of the oceanic tide at Kishu are calculated and are listed in Table 2, 3, 4 and 5, respectively. The vectors of these observed and theoretical values of the main strains are shown in Fig. 2, 3, 4, 5 and 6 . The elements and the main strains of the load tide are calculated for the sea region within 300 km at Osakayama and those within 100 km at Kishu and Suhara.


Fig. 1. Locations of observatories.

Table 1. Characters of the observations and of analysed constants
OSAKAYAMA $\quad 135^{\circ} 51.5^{\prime} \mathrm{E}, \quad 34^{\circ} 59.6^{\prime} \mathrm{N}$


Table 2. Strain elements and main strains for $\mathbf{M}_{2}$-component at Osakayama.

|  | Observed Value |  | Theoretical Value |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Amplitude$\times 10^{-8}$ | Phase | $h_{2}=0.612, l_{2}=0.083$ |  | $h_{2}=0.600, l_{2}=0.200$ |  |
|  |  |  | Amplitude $\times 10^{-8}$ | Phase | Amplitude $\times 10^{-8}$ | Phase |
| $\mathrm{e}_{\text {Tr }}$ | 0.588 | $196.4^{\circ}$ | 0.673 | $180^{\circ}$ | 0.753 | $180^{\circ}$ |
| eөө | 0.612 | 25.7 | 1.356 | 0 | 1.770 | 0 |
| $e_{\phi \phi}$ | 1.109 | 359.0 | 0.511 | 0 | 1.017 | 180 |
|  | 1.080 | 172.7 | 0.490 | 270 | 0.458 | 270 |
| $\mathrm{e}_{\theta \theta}+\mathrm{e}_{\phi \phi}$ | 1.678 | 1.2 | 1.867 | 0 | 0.752 | 0 |
| $\Delta$ | 1.109 | 1.6 | 1.194 | 0 | 0.000 | - |
|  | $\begin{array}{\|c} \hline \text { Amplitude } \\ \times 10^{-8} \\ \hline \end{array}$ | Azimuth | $\begin{array}{\|c} \text { Amplitude } \\ \times 10^{-8} \\ \hline \end{array}$ | Azimuth | $\begin{gathered} \text { Amplitude } \\ \times 10^{-8} \\ \hline \end{gathered}$ | Azimuth |
|  | 1.124 | $121.3^{\circ}$ | 1.356 | $0^{\circ}$ | 1.770 | $0^{\circ}$ |
|  | 0.042 | 31.3 | 0.511 | 90 | -1.017 | 90 |
|  | 0.281 | 102.9 | -0.490 | 45 | -0.229 | 45 |
|  | -0.035 | 12.9 | 0.490 | 135 | 0.229 | 135 |



Fig. 2. Main strains of the observed, theoretical and load-tidal strains for $\mathbf{M}_{2}$-component at Osakayama.

Table 3. Strain elements and main strains for $\mathrm{O}_{1}$-component at Osakayama.

|  | Observed Value |  | Theoretical Value |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Amplitude$\times 10^{-8}$ | Phase | $h_{2}=0.612, l_{2}=0.083$ |  | $h_{2}=0.600, l_{2}=0.200$ |  |
|  |  |  | $\begin{gathered} \text { Amplitude } \\ \times 10^{-8} \end{gathered}$ | Phase | $\underset{\times 10^{-8}}{\substack{\text { Amplitude }}}$ | Phase |
| $\mathrm{e}_{\mathrm{rr}}$ | 0.647 | $214.3{ }^{\circ}$ | 0.279 | $180^{\circ}$ | 0 | - |
| $\mathrm{e}_{\theta \theta}$ | 0.553 | 23.6 | 0.538 | 0 | 0.299 | $180^{\circ}$ |
| $\mathrm{e}_{\phi \phi}$ | 0.698 | 5.0 | 0.666 | 0 | 0.299 | 0 |
| $\mathrm{e}_{\theta \boldsymbol{\phi}}$ | 0.972 | 262.0 | 0.322 | 90 | 0.979 | 90 |
| $\mathrm{e}_{\theta \theta}+\mathrm{e}_{\theta \theta}$ | 1.242 | 13.2 | 1.204 | 0 | 0 | - |
| $\Delta$ | 0.672 | 353.3 | 0.925 | 0 | 0 | - |
|  | $\begin{gathered} \text { Amplitude } \\ \times 10^{-8} \end{gathered}$ | Azimuth | $\begin{gathered} \text { Amplitude } \\ \times 10^{-8} \end{gathered}$ | Azimuth | $\begin{gathered} \text { Amplitude } \\ \times 10^{-8} \\ \hline \end{gathered}$ | Azimuth |
| main $\int \cos$ term $\left\{{ }^{e_{1}}\right.$ | 1.062 | $129.1{ }^{\circ}$ | 0.666 | $90^{\circ}$ | -0.299 | $0^{\circ}$ |
| main $\left\{\right.$ cos term $\left\{\begin{array}{l}e_{2}\end{array}\right.$ | 0.141 | 39.1 | 0.538 | 0 | 0.299 | 90 |
| strain $\left\{\sin\right.$ term $\left\{{ }^{e_{1}}\right.$ | $0.204$ | $141.9$ | 0.161 | 45 | 0.490 | 45 |
| strain $\left(\sin\right.$ term $\left\{\begin{array}{l}e_{1} \\ e_{2}\end{array}\right.$ | 0.078 | 51.9 | -0.161 | 135 | -0.490 | 135 |



Fig. 3. Main strains of the observed, theoretical and load-tidal strains for $\mathrm{O}_{1}$-component at Osakayama.

Table 4. Strain elements and main strains for $\mathrm{M}_{2}$ and $\mathrm{O}_{1}$ components at Kishu.

|  | Observed Value |  |  |  | Loading ( $\mathrm{r}<100 \mathrm{~km}$ ) |  |  |  | Atmospheric Effect <br> S. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\mathrm{M}_{2}$ |  | $\mathrm{O}_{1}$ |  | $\mathrm{M}_{2}$ |  | $\mathrm{O}_{1}$ |  |  |  |
|  | $\begin{aligned} & \text { Amplitude } \\ & 10^{--} \end{aligned}$ | Phase | $\begin{aligned} & \text { Amplitude } \\ & \times 10^{--} \end{aligned}$ | Phase | cos-term | sin-term | cos-term | sin-term | Amplitude $10^{-8}$ | Phase |
| $\mathrm{e}_{1 / \prime}$ | 0.657 | $350.1^{\circ}$ | 0.214 | 3.0 | 32.15 | -11.09 | 27.55 | 0.16 | 0.062 | $112.0^{\circ}$ |
| $\mathrm{e}_{\phi \phi}$ | 0.085 | 124.3 | 0.280 | 20.9 | -32.15 | 11.09 | -27.55 | -0.16 | 0.133 | 112.0 |
| $\mathrm{e}_{\%}$ | 1.070 | 105.9 | 0.388 | 286.3 | -97.02 | 18.98 | 65.82 | -3.20 | 0.024 | 202.0 |
| $e_{\text {r }}$ | - | - | - | - | - | - | - | - | $0.065 \sim 0.097$ | 292.0 |
| $e_{a n}+\mathrm{e}_{\varphi n}$ | 0.600 | 355.8 | 0.488 | 13.2 | 0.00 | 0.00 | 0.00 | 0.00 | 0.195 | 112.0 |
|  | $\begin{gathered} \text { Amplitude } \\ \times 10^{-6} \end{gathered}$ | Azimuth | $\begin{gathered} \text { Amplitude } \\ \times 10^{-8} \end{gathered}$ | Azimuth | Amplitude | Azimuth | Amplitude | Azimuth | $\underset{\times 10^{-8}}{\text { Amplitude }}$ | Azimuth |
| main cos-. $\mathrm{e}_{1}$ | 0.650 | $11.5^{\circ}$ | 0.297 | $56.7{ }^{\circ}$ | 58.20 | $151.8^{\circ}$ | 43.42 | 31.3 | -0.109 | $172.7^{\circ}$ |
| term $\mathrm{e}_{2}$ | -0.051 | 101.5 | 0.178 | 146.7 | -58.20 | 61.8 | -43.42 | 121.3 | $-0.022$ | 82.7 |
|  | 0.501 | 50.1 | 0.247 | 128.3 | 14.60 | 69.7 | 1.68 | 144.6 | 0.123 | 93.9 |
| strain (term $\mathrm{e}_{2}$, | -0.544 | 140.1 | -0.136 | 38.3 | $-14.60$ | 159.7 | -1.68 | 54.6 | 0.057 | 183.9 |



Fig. 4. Main strains of the observed, theoretical and load-tidal strains for $\mathrm{M}_{2}$-component at Kishu.


Fig. 5. Main strain of the observed, theoretical and load-tidal strains for $\mathrm{O}_{1}$-component at Kishu.

Table 5. Strain elements and main strains for $\mathbf{M}_{\mathbf{2}}$-component at Suhara.

|  | Observed Value |  | Loading ( $<100 \mathrm{~km}$ ) |  | $\begin{gathered} \text { Observed Indirect Effect } \\ \left(h_{2}=0.612, l_{2}=0.083\right) \end{gathered}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\underset{\times 10^{-8}}{ }$ | Phase | cos-term | sin-term | $\begin{gathered} \text { Amplitude } \\ \times 10^{-8} \\ \hline \end{gathered}$ | Phase |
| $\mathrm{e}_{\theta \theta}$ | 1.940 | $49.1{ }^{\circ}$ | -2.46 | -6.39 | 1.270 | $90.0^{\circ}$ |
| $\mathrm{e}_{\phi \phi}$ | 1.876 | 36.4 | 2.46 | 6.39 | 1.437 | 50.6 |
| $\mathrm{e}_{\theta \phi}$ | 2.092 | 96.2 | 374.90 | -20.80 | 2.451 | 256.6 |
| $\mathrm{e}_{\theta \theta}+\mathrm{e}_{\phi \phi}$ | 2.780 | 42.8 | - | - |  |  |
|  | $\begin{gathered} \text { Amplitude } \\ \times 10^{-8} \\ \hline \end{gathered}$ | Azimuth | Amplitude | Azimuth | Amplitude $\times 10^{-8}$ | Azimuth |
| main $\left\{\begin{array}{l}\text { cos- } \\ \text { term }\end{array} \begin{array}{l}\mathbf{e}_{1} \\ \mathbf{e}_{2} \\ \text { strain }\end{array}\right.$ | 1.777 | $47.2^{\circ}$ | 187.47 | $45.4{ }^{\circ}$ | 0.996 | $212.0^{\circ}$ |
|  | 1.199 | 137.2 | -187.47 | 135.4 | $-0.084$ | 122.2 |
|  | 2.200 | 42.8 | 12.21 | 105.7 | 2.385 | 86.2 |
|  | 0.183 | 132.8 | -12.21 | 15.7 | -0.002 | 176.2 |



Fig. 6. Main strains of the observed, theoretical and load-tidal strains for $\mathrm{M}_{2}$-component at Suhara.

## 4. Interpretations by means of weighted functions

It is hard to find directly the relations between the observed main strains and the theoretical ones by comparison at first glance of these vectors in Fig. 2, 3, 4 and 5 from the view-point of homogeneous earth crust. Now, we should attempt to find a relationship between the theoretical main strains of the homogeneous earth crust and the observed main strains by means of like a weighted functions. The relationship between these strains is analysed by means of the weighted means in the view-point that the earth crust is anisotropy as followings.
a) Osakayama. From formula (1), (2) and (3) and the Table 2 and 3, we have the weighted functions between the observed tidal main-strains and those of the theoretical values in the case which $h_{2}=0.612$ and $l_{2}=0.083$ as followings, for $M_{2}$-component

$$
\begin{aligned}
& \varphi(\tau)=\frac{1}{2 \pi} \frac{0.831 \cos 2 t+0.123 \sin 2 t}{0.934 \cos 2 t}+\frac{1}{\pi} \frac{1}{(0.423 \cos 2 t)^{2}+(0.245 \sin 2 t)^{2}} \times \\
& \quad[\{0.423 \cos 2 t(-0.279 \cos 2 t+0.142 \sin 2 t)+0.245 \sin 2 t(-0.536 \cos 2 t \\
& \quad+0.069 \sin 2 t)\} \cos 2 \tau+\{-0.245 \sin 2 t(-0.279 \cos 2 t+0.142 \sin 2 t) \\
& \quad+0.423 \cos 2 t(-0.536 \cos 2 t-0.069 \sin 2 t)\} \sin 2 \tau]
\end{aligned}
$$

This result is generally given as the function of the time. In order to obtain the values at the specific times, we separate this weighted function into cosine and sine terms as follows,

$$
\begin{align*}
& \text { cos-term : } \varphi_{\cos (\tau)}=0.1416+0.1548 \cos \left(2 \tau-117.5^{\circ}\right) \text {, } \\
& \text { sin-term }: \varphi_{\sin (\tau)}=\infty \quad+0.1544 \cos \left(2 \tau-305.5^{\circ}\right) \text {, } \\
& =\infty-0.1544 \cos \left(2 \tau-125.5^{\circ}\right) \text {. } \tag{I-2}
\end{align*}
$$

The infinity of the first term in sin-term is caused by the error of the observations, and this term is negligible. The values of the cos-term is equal to the value of $\varphi(\tau)$ at 0 a.m. and p.m., and that of sin-term is equal to the value of $\varphi(\tau)$ at 3 a.m. and p.m., For $0_{1}$-component, we have the weighted function as follows,

$$
\begin{align*}
& \varphi(\tau)=\frac{1}{2 \pi} \frac{0.601 \cos t+0.141 \sin t}{0.602 \cos t}+\frac{1}{\pi} \frac{1}{(0.064 \cos t)^{2}+(0.161 \sin t)^{2}} \times \\
& \quad[\{0.064 \cos t(-0.094 \cos t+0.080 \sin t)+0.161 \sin t(-0.451 \cos t \\
& \quad-0.182 \sin t)\} \cos 2 \tau+0.161 \sin t(-0.094 \cos t+0.080 \sin t) \\
& \quad-0.064 \cos t(-0.451 \cos t-0.182 \sin t)\} \sin 2 \tau] . \tag{I-3}
\end{align*}
$$

This function is separate similarly as follows,

$$
\begin{array}{ll}
\cos \text {-term } & : \quad \varphi_{\cos (\tau)}\left(\tau 0.1589+2.2916 \cos \left(2 \tau-101.7^{\circ}\right),\right. \\
\text { sin-term } & : \quad \varphi_{\sin (\tau)}=\infty-0.3907 \cos \left(2 \tau-156.4^{\circ}\right) . \tag{I-4}
\end{array}
$$

The values of cos-term is equal to the value of $\varphi(\tau)$ at midnight and noon, the value of sin-term is equal to the value of $\varphi(\tau)$ at six a.m..

We can find that these phase angles in these results of $(I-1),(I-2),(I-3)$ and $(I-4)$ are confined in the range from $101.7^{\circ}$ and $156.4^{\circ}$. Therefore, we can estimate that the result shows the crust deformed remarkably in the direction of this range, and it shows that the crust is deformed easily in this direction. This direction coincides with the direction of the axis of Japan Proper.
b) Kishu. Similarly, we have weighted functions for the results at Kishu as follows, for $M_{2}$-component,

$$
\begin{aligned}
& \varphi(\tau)=\frac{1}{2 \pi} \frac{0.300 \cos 2 t-0.022 \sin 2 t}{0.959 \cos 2 t}+\frac{1}{\pi} \frac{1}{(0.417 \cos 2 t)^{2}-(0.244 \sin 2 t)^{2}} \times \\
& {[\{0.417 \cos 2 t(0.348 \cos 2 t-0.092 \sin 2 t)-0.244 \sin 2 t(0.147 \cos 2 t} \\
& -0.515 \sin 2 t)\} \cos 2 \tau+\{-0.244 \sin 2 t(0.348 \cos 2 t-0.092 \sin 2 t) \\
& -0.417 \cos 2 t(0.147 \cos 2 t-0.515 \sin 2 t)\} \sin 2 \tau] \\
& \cos -\text { term }: \varphi_{\cos (\tau)=0.0498+0.291 \cos \left(2 \tau-337.3^{\circ}\right),}^{\sin -\text { term }: \quad 4 \sin (\tau)=\infty+0.682 \cos \left(2 \tau-349.9^{\circ}\right)} .
\end{aligned}
$$

for $0_{1}$-component

$$
\begin{aligned}
& \varphi(\tau)=\frac{1}{2 \pi} \frac{0.237 \cos t+0.056 \sin t}{0.534 \cos t}+\frac{1}{\pi} \frac{1}{(0.122 \cos t)^{2}+(0.203 \sin t)^{2}} \times \\
& \quad[\{-0.122 \cos t(-0.024 \cos t-0.045 \sin t)+0.203 \sin t(0.055 \cos t \\
& -0.186 \sin t)\} \cos 2 \tau-\{0.203 \sin t(-0.024 \cos t+0.045 \sin t) \\
& +0.122 \cos t(0.055 \cos t-0.186 \sin t)\} \sin 2 \tau], \\
& \quad \cos -\text { term } \quad: \varphi_{\cos (\tau)=0.140+0.157 \cos \left(2 \tau-66.6^{\circ}\right),} \quad \operatorname{sin-term}: \varphi_{\sin }(\tau)=\infty-0.300 \cos \left(2 \tau-344.4^{\circ}\right) .
\end{aligned}
$$

According to this result, these phase angles are confined within the range from $337.3^{\circ}$ to $349.9^{\circ}$ except $66.6^{\circ}$ at Kishu. We might estimate that the crust is deformable easy in the direction in this range from $337.3^{\circ}$ to $349.9^{\circ}$ in this district. Namely, it shows that the crust at Kishu which is in the eastern shoulder of the Kii Peninsular has been generated in the direction of the south-south-east.
c) Suhara. Similarly, we have the weighted function at Suhara as follows,

$$
\begin{align*}
& \varphi(\tau)=\frac{1}{2 \pi} \frac{1.513 \cos 2 t+1.192 \sin 2 t}{0.322 \cos 2 t}+\frac{1}{\pi} \frac{1}{(0.4352 \cos 2 t)^{2}+(0.1855 \sin 2 t)^{2}} \times \\
& \quad[\{0.435 \cos 2 t(-0.0469 \cos 2 t+0.0787 \sin 2 t)-0.1855 \sin 2 t(-0.1275 \cos 2 t \\
& +1.0380 \sin 2 t)\} \cos 2 \tau+\{-0.1855 \sin 2 t(-0.0469 \cos 2 t+0.0787 \sin 2 t) \\
& -0.4352 \cos 2 t(-0.1275 \cos 2 t+1.0380 \sin 2 t)\} \sin 2 \tau] \\
& \quad \cos \text {-term }: \varphi_{\cos (\tau)=0.233-0.099 \cos \left(2 \tau-290.2^{\circ}\right),} \begin{array}{l}
\sin \text {-term } \quad: \quad \varphi_{\sin (\tau)}(\tau)+1.786 \cos \left(2 \tau-184.3^{\circ}\right) .
\end{array} \quad \ldots \text { (III) }
\end{align*}
$$

The author expected that the direct effect was negligible for the load tide because the Suhara Observatory was at coast. But, according to our observations, the load tide is not so much larger than the direct effect.

Now, the resultant effects $e_{\theta \theta^{\prime}}, e_{\phi \phi, \xi^{\prime}}$ and $e_{\phi \phi}{ }^{\prime}$ which are differences between the observed values and the theoretical direct effect in the case of $h_{2}=0.612$ and $l_{2}=0.083$ are calculated as follows,

$$
\left.\begin{array}{l}
e_{\theta \theta^{\prime}}=(-0.0010 \cos 2 t+1.2701 \sin 2 t) \times 10^{-8}  \tag{III'}\\
e_{\phi \phi}{ }^{\prime}=\left(\begin{array}{r}
0.9130 \cos 2 t+1.1132 \sin 2 t) \times 10^{-8}, \\
e_{\theta \phi}^{\prime}
\end{array}\right\}(-0.5766 \cos 2 t+2.3819 \sin 2 t) \times 10^{-8} .
\end{array}\right\}
$$

And the weighted function $\varphi(\tau)$ between the theoretical load-tide and the resultants are calculated as follows

$$
\begin{aligned}
& \varphi(\tau)=\frac{1}{\pi} \cdot \frac{1}{(-2.46 \cos 2 t-6.39 \sin 2 t)^{2}+(187.45 \cos 2 t-10.40 \sin 2 t)^{2}} \times \\
& \quad[\{(187.45 \cos 2 t-10.40 \sin 2 t)(-0.4570 \cos 2 t+0.0785 \sin 2 t) \\
& \quad-(-2.46 \cos 2 t-6.39 \sin 2 t)(-0.2883 \cos 2 t+1.1910 \sin 2 t)\} \cos 2 \tau \\
& +(187.45 \cos 2 t-10.40 \sin 2 t)(-0.4570 \cos 2 t+0.0785 \sin 2 t) \\
& -(-2.46 \cos 2 t-6.39 \sin 2 t)(-0.2883 \cos 2 t+1.1910 \sin t)\} \sin 2 \tau] \\
& \quad \cos -\text { term } \quad: \varphi_{\cos }(\tau)=\infty+0.001948 \cos \left(2 t-311.1^{\circ}\right) \\
& \quad \sin -\text { term } \quad: \varphi \sin (\tau)=\infty+0.2179 \cos \left(2 t-302.6^{\circ}\right)
\end{aligned}
$$

According to this result, the phase angles of $\varphi_{\cos (\tau)}$ and $\varphi_{\sin }(\tau)$ are $311.1^{\circ}$ and $312.6^{\circ}$, respectively, and these almost equal each other. We can estimate that the crust around Suhara is deformed with the deflection of about $45^{\circ}$ anti-clockwisely for the stress caused by the oceanic tide. The present author thinks that this deflection is caused by the geological constructions which consists of metamorphic rocks over the large area.

As these summaries, the tidal strains in Kinki District are much deflected by the constructions of the arc island, peninsular and the median line of the geological constructions.

And the present author suceeded in the expressions of the anomalous strain by means of the weighted functions.

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