

# Movable singularity of generalized Emden equation via Birkhoff reduction

By

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## Abstract

This paper studies the movable singularity of the generalized Emden-Fowler equation. For a certain nonlinear term we shall show the existence of infinitely many movable singularities in the complex plane by virtue of the reduction similar to Birkhoff's normal form theory.

## § 1. Motivation and the result

In this paper we shall study the movable singularity of generalized Emden-Fowler equation

$$(1.1) \quad \frac{d^2u}{dt^2} + \frac{n-1}{t} \frac{du}{dt} + u^\ell = 0, \quad t \in \mathbb{C},$$

where  $n \neq 2, \ell \geq 2$  are integers. By the movable singularity we mean the singularity which depends on the respective solution and that does not appear in the coefficients of the equation. There are many works as to the movable singularity of nonlinear ordinary differential equation. We cite [1] and the references therein. The existence of movable branching singularity of (1.1) was shown in [1] by the argument closely related with Painlevé's theorem. Recently, detailed analysis on the singularity was also made in terms of phase space analysis. (cf. [4]). The object of the present paper is to give an alternative proof of the existence of a movable singularity by virtue of the reduction argument similar to the Birkhoff normal form theory of a Hamiltonian system. The advantage of our method is that one can show some structure of the movable singularity

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in terms of Jacobi's elliptic function. Indeed, we write (1.1) in the Hamiltonian system and we transform it to the autonomous form by some analytic transformation in certain domain. By virtue of Jacobi's elliptic function we shall show the existence of infinitely many movable singularities. We shall prove

**Theorem 1.1.** *Assume that  $\ell = 3$  in (1.1). Then (1.1) has an infinitely many movable singularities.*

## § 2. Normalization via Birkhoff reduction

We shall transform (1.1) to a simpler form. For this purpose we write (1.1) in a Hamiltonian system. Set  $\ell = 3$ ,  $t = e^z$  and define  $v = e^{(2-n)z/2}u$ . Then, by (1.1) we have

$$(2.1) \quad u_{zz} - \tilde{c}_0 u + e^{z\beta} u^3 = 0,$$

where  $\tilde{c}_0 = (n-2)^2/4$  and  $\beta = 4-n$ . Next, by putting  $q := u$  and  $p := u_z$ , we write (2.1) in the system of equations for  $q$  and  $p$ . Indeed, the Hamiltonian function is given by

$$\frac{1}{2}(p^2 - \tilde{c}_0 q^2) + \frac{e^{z\beta}}{4} q^4.$$

By replacing  $q$  and  $p$  with constant times of respective variable, if necessary, and by introducing new unknown quantities  $q_2 := q+p$  and  $p_2 := q-p$  we can write the system in the Hamiltonian system with Hamiltonian

$$(2.2) \quad H = \lambda q_2 p_2 + c e^{z\beta} (q_2 + p_2)^4,$$

where  $\lambda \neq 0$  and  $c$  are certain constants. One may assume  $\lambda = 1$  without loss of generality.

We set  $q_1 = z$  and  $\tilde{c}(z) = 4c e^{\beta z}$  in (2.2). Consider the autonomous system  $p_1 + H$  with  $H := q_2 p_2 + c e^{z\beta} (q_2 + p_2)^4$ . The resonance term appears from the nonlinear part. Define

$$(2.3) \quad a \equiv a(q_2 p_2) := 6c q_2^2 p_2^2.$$

We shall find a formal Birkhoff transformation which brings the nonlinear part of  $H$  to the one  $a(q_2 p_2) + c(q_2^4 + p_2^4)$ . In fact, we have

**Theorem 2.1.** *There exists a formal symplectic transformation which transforms  $p_1 + H$  to  $p_1 + q_2 p_2 + a(q_2 p_2) + c(q_2^4 + p_2^4)$ .*

*Proof.* The proof is essentially Birkhoff's reduction. For the sake of completeness we give the proof. By the general theory of the transformation of a vector field, we know that the vector field  $X_0 + R$  is transformed to  $X_0 + S$  by certain coordinate change generated by  $Y$ , if we determine  $Y$  by the homology equation (adjoint equation)  $[X_0, Y] = R - S$ , where  $[X_0, Y]$  denotes the Lie bracket and  $S$  is a certain term containing the resonance. We shall solve the homology equation  $[X_0, Y] = R - S$  formally in terms of the infinite composition of symplectic transformations. Because we cannot show the convergence of the infinite compositions, the transformation is formal. As to the solvability and properties of the adjoint equation in the Hamiltonian setting we refer to Lemma 6 in [3]. Let  $\chi_g$  denotes the Hamiltonian vector field with the Hamiltonian function  $g$  with respect to a standard symplectic structure. We denote by  $\{\cdot, \cdot\}$  the Poisson bracket.

Let  $x = (\tilde{q}_1, \tilde{p}_1, \tilde{q}_2, \tilde{p}_2)$  and  $y = (q_1, p_1, q_2, p_2)$  be the original and the transformed coordinates, respectively. For simplicity we sometimes write  $y = (y_1, \dots, y_4)$ . We consider the transformation  $x = u(y)$  for some  $u = (u_1, \dots, u_4)$ . For  $\phi \equiv \phi(q_1, q_2, p_2)$ , define

$$(2.4) \quad X_0 := \chi_{p_1+q_2p_2} \quad Y := \chi_\phi = \sum_j u_j \frac{\partial}{\partial y_j}.$$

Then the component of  $X_0$  on the basis  $\partial/\partial q_1, \partial/\partial p_1, \partial/\partial q_2, \partial/\partial p_2$  with this order is given by  $\Lambda(y) = (1, 0, q_2, -p_2)$ . Set

$$(2.5) \quad f(x) := \tilde{c}(\tilde{q}_1)(\tilde{q}_2 + \tilde{p}_2)^4/4,$$

and define

$$(2.6) \quad R := \chi_f, \quad S := \chi_{a(q_2p_2)+c(q_2^4+p_2^4)}.$$

Write  $R = r(x) \frac{\partial}{\partial x}$  and  $S = s(y) \frac{\partial}{\partial y}$ . Note that the first and the second components of  $s(y)$  vanish and  $r(x)$  is independent of  $x_2$ . We have

$$(2.7) \quad \chi_{\{p_1+q_2p_2, \phi\}} = [X_0, Y] = X_0Y - YX_0.$$

Note that the component of the right-hand side is given by  $\Lambda(y)\nabla u - u\nabla\Lambda(y)$ . Then, by simple computation the equation  $[X_0, Y] = R - S$  is written as

$$(2.8) \quad \Lambda(y)\nabla u - u\nabla\Lambda(y) = r(u) - s(y)\nabla u.$$

If we set  $u_1 = v_1, u_2 = v_2, u_j(y) = y_j + v_j(y)$ , then we see that (2.8) is written in the form

$$(2.9) \quad \Lambda(y)\nabla v - v\nabla\Lambda(y) = r(v_1, y_j + v_j) - s(y) - s(y)\nabla v.$$

Clearly, if we can solve (2.8) or (2.9) formally, then we have our theorem. Approximate the right-hand side of (2.8) with  $r(y) - s(y)$ . Consider the linear equation of  $\phi$

$$(2.10) \quad \{p_1 + q_2 p_2, \phi\} = f(y) - a(q_2 p_2) - c(q_2^4 + p_2^4).$$

The right-hand side is homogeneous degree of four. Consider the monomial of the form  $e^{\beta q_1} K q_2^i p_2^j$  for some constants  $\beta$  and  $K$  in the right-hand side. Set  $\phi = e^{\beta q_1} \psi(q_2, p_2)$ . Clearly, if  $\beta + i - j \neq 0$ , then one can determine  $\phi$ . If otherwise, then we set  $\phi = q_1 e^{\beta q_1} \psi(q_2, p_2)$ . In general, we consider  $\alpha(q_1) \beta(q_2, p_2)$ , where  $\beta(q_2, p_2)$  is a monomial. We have, for  $\phi = \omega(q_1) \psi(q_2, p_2)$

$$\{p_1 + q_2 p_2, \phi\} = \{p_1, \omega\} \psi + \{q_2 p_2, \psi\} \omega = \alpha(q_1) \beta(q_2, p_2).$$

Clearly,  $\{q_2 p_2, \psi\}$  can be chosen as the constant times of  $\beta(q_2, p_2)$ . Hence  $\omega$  satisfies the equation like  $\omega' + c_1 \omega = c_2 \alpha$  for some constants  $c_1$  and  $c_2$ . Therefore one can solve (2.10).

One can verify that  $\phi$  gives the terms of homogeneous degree 4 of  $v$  in (2.9). Set  $u = y + v$ . Then, the change of variables  $x = u(y)$  eliminate the terms of degree 4 in  $f$  except for those which remain in  $S$ . After the change of variables we make the same argument in such a way that we eliminate the terms of homogeneous degree 5 in the right-hand side of (2.9). We repeat the same argument. One can delete every homogeneous term in the right-hand side of (2.9). Finally, by the infinite compositions of transformations like  $u = y + v$  one can solve (2.9) in category of formal power series.  $\square$

Next we give a meaning to a certain normal form obtained in Theorem 2.1. First we look for the alternative expression of the homology equation. In fact, we have

**Lemma 2.2.** *Suppose that  $u$  satisfy*

$$(2.11) \quad \Lambda(y) \nabla u + s(y) \nabla u = r(u) + \Lambda(u).$$

*Then, the transformation  $x = u(y)$  maps the vector field  $(\Lambda(x) + r(x)) \frac{\partial}{\partial x}$  to  $(\Lambda(y) + s(y)) \frac{\partial}{\partial y}$ .*

*Proof.*

$$(\Lambda(x) + r(x)) \frac{\partial}{\partial x} = (\Lambda(u) + r(u)) (\nabla u)^{-1} \frac{\partial}{\partial y} = (\Lambda(y) + s(y)) \frac{\partial}{\partial y}.$$

$\square$

We remark that  $\Lambda(u)$  in (2.11) and  $u \nabla \Lambda(y)$  in (2.8) is identical except for the term 1 in the first component, which causes a minor change of variables in the original

equation. We shall simplify (2.11). We first consider the third equation of (2.11). Set  $w = u_3$ ,  $u = (u_1, \dots, u_4)$ . Then we have

$$(2.12) \quad \frac{\partial w}{\partial q_1} + \delta w + q_2 p_2 \delta + p_2^3 \frac{\partial w}{\partial q_2} - q_2^3 \frac{\partial w}{\partial p_2} - w - \mathcal{R}(u) = 0$$

where  $\delta = q_2 \frac{\partial}{\partial q_2} - p_2 \frac{\partial}{\partial p_2}$  and  $\mathcal{R}(u) = \tilde{c}(u_1)(u_3 + u_4)^3$ . Because we consider a singular solution we shall solve (2.12) near  $q_2 = \infty$ ,  $p_2 = \infty$ . Hence we replace  $q_2 \mapsto q_2^{-1}$ ,  $\frac{\partial}{\partial q_2} \mapsto -q_2^2 \frac{\partial}{\partial q_2}$  and so on in (2.12) and we multiply the equation with  $q_2^3 p_2^3$ . Then we obtain

$$(2.13) \quad q_2^3 p_2^3 \frac{\partial w}{\partial q_1} - q_2^3 p_2^3 \delta w - q_2^2 p_2^2 \delta w - q_2^5 \frac{\partial w}{\partial q_2} + p_2^5 \frac{\partial w}{\partial p_2} - q_2^3 p_2^3 w - q_2^3 p_2^3 \mathcal{R}(u) = 0.$$

We shall look for  $w = w(q_1, q_2 + p_2)$ . Set  $r = q_2 + p_2$  and define

$$(2.14) \quad q_2 = \alpha \zeta, \quad p_2 = \eta \zeta,$$

where  $\alpha + \eta \neq 0$  and  $\zeta$  is a complex parameter. Then we have  $\delta w = (\alpha - \eta) \zeta \frac{\partial w}{\partial r}$  and

$$-q_2^5 \frac{\partial w}{\partial q_2} + p_2^5 \frac{\partial w}{\partial p_2} = (-\alpha^5 + \eta^5) \zeta^5 \frac{\partial w}{\partial r}.$$

Hence, by (2.13) we have

$$(2.15) \quad \zeta \frac{\partial w}{\partial q_1} + (\zeta^2(\eta - \alpha) + \alpha^{-1} - \eta^{-1}) \frac{\partial w}{\partial r} + (-\alpha^2 \eta^{-3} + \alpha^{-3} \eta^2) \frac{\partial w}{\partial r} - \zeta w - \zeta \mathcal{R}(u) = 0.$$

Suppose that  $\alpha^{-1} - \eta^{-1} - \alpha^2 \eta^{-3} + \alpha^{-3} \eta^2 \neq 0$ . This is equivalent to  $(\eta^3 - \alpha^3)(\eta^2 + \alpha^2) \neq 0$ . We assume

$$(2.16) \quad \eta^3 - \alpha^3 \neq 0, \quad \eta^2 + \alpha^2 \neq 0, \quad \eta + \alpha \neq 0.$$

We make the change of variables  $q_1 \mapsto \lambda q_1$  and  $r \mapsto rK$  for some  $\lambda \neq 0$  and  $K \neq 0$  such that (2.15) is transformed to the following

$$(2.17) \quad r \frac{\partial w}{\partial q_1} + (r^2 + 1) \frac{\partial w}{\partial r} - \epsilon_0 r w - \epsilon_0 r \mathcal{R}(u) = 0,$$

where  $\epsilon_0$  is some constant. Similarly, one can verify that  $u_4$  satisfies (2.17) with  $-\epsilon_0 r w$  replaced by  $\epsilon_0 r w$ .

Next we consider the equation for  $u_1$  in (2.11), which is given by the first column of (2.11). Because the first component of  $r(u)$  vanishes, the equation for  $w := u_1$  is exactly the equation (2.17) with  $-\epsilon_0 r w - \epsilon_0 r \mathcal{R}(u)$  replaced by  $-\epsilon_0 r$ . One can easily verify that it has a solution

$$(2.18) \quad \tilde{q}_1 \equiv u_1 = q_1 - \frac{1 - \epsilon_0}{2} \log(1 + r^2).$$

On the other hand, the equation for  $u_2$  has the same form as (2.17), where the nonlinear term  $\mathcal{R}(u)$  does not contain  $u_2$ . Hence the equation is linear. One can integrate it by the method which we show in the following if we determine  $u_3$  and  $u_4$ .

Insert  $u_1$  into  $\mathcal{R}(u) = \tilde{c}(u_1)(u_3 + u_4)^3$  and consider the system of equations for  $u_3$  and  $u_4$  in (2.17). We note that one can delete  $\epsilon_0 r w$  by introducing the unknown quantity  $w_3 = w e^{q_1 \epsilon_0}$ . Then  $\tilde{c}$  in  $\mathcal{R}$  is multiplied by some power of  $e^{q_1 \epsilon_0}$ , which we denote by  $\mathcal{R}_3$ . Defining  $w_4 = w e^{-q_1 \epsilon_0}$  and  $\mathcal{R}_4$  in the same way as  $\mathcal{R}_3$  we obtain

$$(2.19) \quad r \frac{\partial w_3}{\partial q_1} + (1 + r^2) \frac{\partial w_3}{\partial r} = \epsilon_0 r \mathcal{R}_3(w_3, w_4).$$

$$(2.20) \quad r \frac{\partial w_4}{\partial q_1} + (1 + r^2) \frac{\partial w_4}{\partial r} = \epsilon_0 r \mathcal{R}_4(w_3, w_4).$$

We shall solve the reduced homology equation (2.19) and (2.20) by iteration. Consider

$$(2.21) \quad r \frac{\partial W}{\partial q_1} + (1 + r^2) \frac{\partial W}{\partial r} = f(r, q_1)$$

where  $W = (w_3, w_4)$  and  $f(r, q_1)$  is a given vector function.

The equation (2.21) can be solved by the method of characteristics. Consider

$$(2.22) \quad \frac{dq_1}{r} = \frac{dr}{1 + r^2} = d\sigma.$$

By integration we have  $\tan^{-1} r = \sigma + \tan^{-1} r^0$  and  $q_1 = q_1^0 - \log \cos \sigma$ . For given  $r$  and  $q_1$  we set  $r^0 = 0$  and define  $\sigma_0$  and  $q_1^0$  by

$$(2.23) \quad \tan^{-1} r = \sigma_0, \quad q_1 = q_1^0 - \log \cos \sigma_0.$$

Then the solution of (2.21) is given by

$$(2.24) \quad W(r, q_1) = Pf := \int_0^{\sigma_0} f(\tan \sigma, q_1^0 - \log \cos \sigma) d\sigma.$$

We shall show that in the integrand of (2.24) the function  $\tilde{c}(\tilde{q}_1)$  is bounded, where  $\tilde{q}_1 = q_1 - \frac{1-\epsilon_0}{2} \log(1+r^2)$ . Because  $\tilde{c}(\tilde{q}_1)$  is a constant times of  $e^{\beta \tilde{q}_1}$  we consider the latter function. By using  $q_1^0 - \log \cos \sigma = q_1 + \log \cos \sigma_0 - \log \cos \sigma$  and  $r = \tan \sigma$  we have

$$(2.25) \quad e^{\beta \tilde{q}_1} = e^{\beta q_1} (1 + r^2)^{-\beta/2 + \epsilon_0/2} = e^{\beta q_1} (\cos \sigma_0)^\beta (\cos \sigma)^{-\epsilon_0}.$$

Hence the integrand is bounded in  $\sigma$ .

Define

$$(2.26) \quad \Omega_0 := \{(q_1, r) \mid |q_1| < \eta_0, |r| < \eta_0\}.$$

We shall solve (2.19) and (2.20) in the set of functions analytic in  $\Omega_0$  and continuous up to the boundary. Set  $K := \overline{\Omega_0}$  and define

$$(2.27) \quad \|W\| := \|w_3\| + \|w_4\|, \quad \|w_j\| := \sup_K |w_j(r, q_1)|.$$

Define  $Q \equiv Q(W) := \epsilon_0 r(\mathcal{R}_3, \mathcal{R}_4)$ . We define the approximate sequence  $\{W_n\}_n$  by  $W_{-1} = 0$ ,  $W_0 = (q_1 - \frac{1}{2} \log(1 + r^2), q_1 - \frac{1}{2} \log(1 + r^2))$  and

$$(2.28) \quad W_n = W_{n-1} + PQ(W_{n-1}) - PQ(W_{n-2}), \quad n = 1, 2, \dots$$

One can easily see that the sequence is well defined on  $\Omega_0$ . If the limit,  $W := \lim W_n = W_0 + \sum_{n=1}^{\infty} (W_n - W_{n-1})$  exists, then we have

$$(2.29) \quad W - W_0 = P \left( \sum_{n=1}^{\infty} (Q(W_{n-1}) - Q(W_{n-2})) \right) = P(\lim Q(W_n)) = PQ(W).$$

By the definition of  $W_0$ ,  $W$  satisfies (2.21).

We shall show an a priori estimate. Given an  $\epsilon > 0$ . Take  $C_1 > 0$  such that  $|e^{\beta \tilde{q}_1}| \leq C_1$  on  $\Omega_0$ . If  $\eta_0$  is sufficiently small, then we have  $\|W_0\| \leq \epsilon/2$ . Consider  $W_1 - W_0 = PQ(W_0)$ . We have  $|r| < \epsilon$  if  $\eta_0$  is sufficiently small. Because  $\tilde{c}(\tilde{q}_1)$  in the integrand is bounded, we get, from the formula of  $P$  that  $\|W_1 - W_0\| < \tilde{K}\epsilon^2$  for some  $\tilde{K} > 0$  depending only on the equation. Take  $\epsilon$  such that  $\tilde{K}\epsilon \leq 1/4$ . Then we have  $\|W_1\| \leq (1/2 + 1/4)\epsilon < \epsilon$ .

Next, we shall estimate  $\|W_2 - W_0\|$ . Writing  $W_2 - W_0 = W_2 - W_1 + W_1 - W_0$  we consider  $\|W_2 - W_1\|$ , where

$$(2.30) \quad W_2 - W_1 = P(Q(W_1) - Q(W_0)).$$

By the definition of  $Q$  we see that  $Q(W_1) - Q(W_0)$  is estimated by the constant times of  $\|W_1 - W_0\|(\|W_1\|^2 + \|W_1\|\|W_0\| + \|W_0\|^2)$ . By the estimate of  $W_1$  the last term is bounded by  $3\epsilon^2\|W_1 - W_0\| \leq 3\epsilon^3/4$ . Hence, by (2.30) we have  $\|W_2 - W_1\| \leq \epsilon^3 K$  for some  $K$  depending only on the equation. Therefore we have

$$(2.31) \quad \|W_2 - W_0\| \leq \epsilon/4 + \epsilon^3 K \leq \epsilon(1/4 + \epsilon^2 K).$$

If  $\epsilon^2 K \leq 1/8$ , then we have  $\|W_2 - W_0\| < \epsilon/2$ . Especially, we have  $\|W_2\| \leq \epsilon$ .

We proceed in the same way for  $W_3 - W_0 = W_3 - W_2 + W_2 - W_1 + W_1 - W_0$ . We can show that  $\|W_3 - W_0\| \leq 2^{-1}\epsilon(1/2 + 1/4 + 1/8) < \epsilon/2$ . In general we have  $\|W_n - W_0\| \leq \epsilon/2$  for all  $n$ . Hence we obtain the a priori estimate

$$(2.32) \quad \|W_n\| \leq \epsilon, \quad n = 0, 1, \dots$$

By the apriori estimate and (2.28) we have

$$(2.33) \quad \|W_n - W_{n-1}\| \leq 3\epsilon^2 \|W_{n-1} - W_{n-2}\| \leq \frac{1}{2} \|W_{n-1} - W_{n-2}\|$$

if  $3\epsilon^2 \leq 1/2$ . This proves the convergence. Therefore we have proved

**Theorem 2.3.** *Suppose that  $\alpha$  and  $\eta$  satisfy (2.16). Then there exists an  $\eta_0 > 0$  such that if  $p_1$  is in some neighborhood of the origin and  $(q_1, q_2, p_2)$  is given by (2.14) with  $q_2$  and  $p_2$  replaced by  $q_2^{-1}$  and  $p_2^{-1}$ , respectively, and  $\zeta = (\alpha + \eta)^{-1}r$ ,  $(q_1, r) \in \Omega_0$ , then the vector field  $(\Lambda(x) + r(x))\frac{\partial}{\partial x}$  is transformed to  $(\Lambda(y) + s(y))\frac{\partial}{\partial y}$  by an analytic change of coordinates.*

### § 3. Movable Singularity

In this section we shall study the movable singularity of solutions of (1.1) with  $\ell = 3$ . In view of Theorem 2.3 we consider the reduced Hamiltonian  $p_1 + q_2 p_2 + c q_2^2 p_2^2 + \tilde{c}(q_2^4 + p_2^4)$ , where  $c \neq 0$  and  $\tilde{c} \neq 0$  are constants. By setting  $q_2 = q$  and  $p_2 = p$  we consider the Hamiltonian

$$(3.1) \quad \tilde{H} := qp + \frac{\epsilon}{2} q^2 p^2 - \frac{\eta}{8} (q^2 - p^2)^2,$$

where  $\epsilon$  and  $\eta \neq 0$  are constants. Because  $\tilde{c}$  can be chosen arbitrarily by changing the transformation in the proof of Theorem 2.3, we may assume  $\epsilon \neq 0$ ,  $\epsilon + \eta \neq 0$  without loss of generality. Suppose that  $(q, p)$  is the solution of the Hamiltonian system for  $\tilde{H}$ . Then there exists a constant  $C_2$  such that  $\tilde{H} = C_2$ . Define

$$(3.2) \quad \zeta = \frac{q+p}{2}, \quad \xi = \frac{q-p}{2i}.$$

Then we have

$$(3.3) \quad \begin{aligned} C_2 = \tilde{H} &= (\zeta^2 + \xi^2) + \frac{\epsilon}{2} (\zeta^2 + \xi^2)^2 + 2\eta \zeta^2 \xi^2 \\ &= \frac{\epsilon + \eta}{2} \left( \zeta^2 + \xi^2 + \frac{1}{\epsilon + \eta} \right)^2 - \frac{1}{2(\epsilon + \eta)} - \frac{\eta}{2} (\zeta^2 - \xi^2)^2. \end{aligned}$$

Hence we have

$$(3.4) \quad 1 = \frac{(\epsilon + \eta)^2}{A} \left( \zeta^2 + \xi^2 + \frac{1}{\epsilon + \eta} \right)^2 - \frac{\eta(\epsilon + \eta)}{A} (\zeta^2 - \xi^2)^2$$

where  $A = 1 + 2C_2(\epsilon + \eta)$ . We determine  $\theta = \theta(z)$  such that

$$(3.5) \quad \sin^2 \theta = \frac{(\epsilon + \eta)^2}{A} \left( \zeta^2 + \xi^2 + \frac{1}{\epsilon + \eta} \right)^2, \quad \cos^2 \theta = -\frac{\eta(\epsilon + \eta)}{A} (\zeta^2 - \xi^2)^2.$$



Then, by (3.5) and simple computations we have

$$(3.6) \quad \zeta \equiv \zeta(z) = \sqrt{\frac{\sqrt{\frac{A(\epsilon+\eta)}{-\eta}} \cos \theta + \sqrt{A} \sin \theta - 1}{2(\epsilon + \eta)}},$$

$$(3.7) \quad \xi \equiv \xi(z) = \sqrt{\frac{-\sqrt{\frac{A(\epsilon+\eta)}{-\eta}} \cos \theta + \sqrt{A} \sin \theta - 1}{2(\epsilon + \eta)}}.$$

Set  $X(z) = \sin \theta(z) + \eta\epsilon^{-1}/\sqrt{A}$  and define

$$(3.8) \quad \mathcal{A} = \sqrt{\mathcal{E} + \frac{i}{2}\sqrt{\mathcal{F}}}, \quad \mathcal{B} = \sqrt{\mathcal{E} - \frac{i}{2}\sqrt{\mathcal{F}}},$$

where

$$(3.9) \quad \mathcal{E} = \frac{1}{2(\epsilon + \eta)} \left( \sqrt{A}X(z) - \eta\epsilon^{-1} - 1 \right),$$

$$(3.10) \quad \mathcal{F} = \frac{A}{\eta(\epsilon + \eta)} \left( 1 - (X(z) - \eta\epsilon^{-1}/\sqrt{A})^2 \right).$$

Then we see that  $\zeta = \mathcal{A}$  and  $\xi = \mathcal{B}$ . Therefore, by (3.2) we obtain

$$(3.11) \quad q(z) = \mathcal{A} + i\mathcal{B}, \quad p(z) = \mathcal{A} - i\mathcal{B}.$$

We will show the following lemma.

**Lemma 3.1.**  *$X(z)$  is an elliptic function.*

*Proof.* By (3.1) the Hamiltonian equation with Hamiltonian  $\tilde{H}$  is given by

$$\dot{q} = q + \epsilon q^2 p + \frac{\eta}{2} p(q^2 - p^2), \quad \dot{p} = -p - \epsilon q p^2 + \frac{\eta}{2} q(q^2 - p^2).$$

In terms of (3.2), these equations are written as

$$(3.12) \quad \dot{\zeta} = i\xi + i\epsilon\xi(\zeta^2 + \xi^2) + 2i\eta\zeta^2\xi, \quad \dot{\xi} = -i\zeta - i\epsilon\zeta(\zeta^2 + \xi^2) - 2i\eta\zeta\xi^2.$$

In view of (3.2) we have

$$(3.13) \quad \frac{d}{dz}(\zeta^2 + \xi^2) = 4\eta i\zeta\xi(\zeta^2 - \xi^2).$$

By (3.6) and (3.7) we have

$$(3.14) \quad \zeta^2 - \xi^2 = \frac{\sqrt{\frac{A(\epsilon+\eta)}{-\eta}} \cos \theta}{\epsilon + \eta}.$$

$$\begin{aligned}
(3.15) \quad \zeta^2 \xi^2 &= \frac{(\sqrt{A} \sin \theta - 1)^2 + \eta^{-1} A(\epsilon + \eta) \cos^2 \theta}{4(\epsilon + \eta)^2} \\
&= \frac{1}{4(\epsilon + \eta)^2} \left( -\eta^{-1} A \epsilon \sin^2 \theta - 2\sqrt{A} \sin \theta + 1 + \eta^{-1} A(\epsilon + \eta) \right) \\
&= \frac{1}{4(\epsilon + \eta)^2} \left( -\eta^{-1} A \epsilon (\sin \theta + \eta \epsilon^{-1} A^{-1/2})^2 + \eta \epsilon^{-1} + 1 + \eta^{-1} A(\epsilon + \eta) \right) \\
&= \frac{1}{4(\epsilon + \eta)^2} \left( -\eta^{-1} A \epsilon X^2 + \eta \epsilon^{-1} + 1 + \eta^{-1} A(\epsilon + \eta) \right).
\end{aligned}$$

$$(3.16) \quad \frac{d}{dz} (\zeta^2 + \xi^2) = \frac{d\theta}{dz} \frac{d}{d\theta} \left( \frac{\sqrt{A} \sin \theta - 1}{\epsilon + \eta} \right) = \frac{\sqrt{A} \cos \theta}{\epsilon + \eta} \frac{d\theta}{dz} = \frac{\sqrt{A}}{\epsilon + \eta} \frac{dX}{dz}.$$

By (3.13), (3.14), (3.15) and (3.16) we have

$$\begin{aligned}
(3.17) \quad \frac{dX}{dz} &= \sqrt{\frac{\eta}{\epsilon + \eta}} \sqrt{1 - (X - \eta \epsilon^{-1} A^{-1/2})^2} \\
&\quad \times \sqrt{-\eta^{-1} A \epsilon X^2 + \eta \epsilon^{-1} + 1 + \eta^{-1} (\epsilon + \eta) A}.
\end{aligned}$$

Hence,  $(\frac{dX}{dz})^2$  is at most the fourth order polynomial of  $X$ . By the general theory of the elliptic function we see that  $X(z)$  is Jacobi's elliptic function. (See [2] and [5]).  $\square$

*Proof of Theorem 1.1.* First we recall that  $qp = \zeta^2 + \xi^2 = C_0 X(z) + C_1$  for some constants  $C_0 \neq 0$  and  $C_1$  by what we have proved in the above, where  $X(z)$  is the elliptic function. In view of representation formula of  $q$  and  $p$ ,  $q$  and  $p$  are singular at the poles of the elliptic function  $X(z)$ . It may occur that another branch point appears for  $p$  or  $q$ . One can easily verify that the singular point is the movable singularity because it does not appear in the coefficients of the equations of  $q$  and  $p$ .

In order to show that the generalized Emden equation has movable singularities, we transform the equation to a normal form as in Theorem 2.3 in some domain near the infinity. Indeed, the assumption (2.16) is the condition on  $q_2$  and  $p_2$ , which can be satisfied by changing parameters in view of the expression of singular solution in §3.

Recall that the elliptic function has a pole on each parallelogram. By suitable choice of parameters of the transformed equation, the image of the fundamental parallelogram lies in the domain on which the normalizing transformation is defined. Hence we have a singular solution of the generalized Emden equation which is parametrized by the elliptic function. Finally, by (2.18) one can change the independent variable of the transformed equation to that of the generalized Emden equation, and one obtains a singular solution of the generalized Emden equation. There exist infinitely many movable singularities in view of the periodicity of the elliptic function.  $\square$

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